

## REVISION 1

### Relaxation effects and re-entrant spin glass behaviour at low temperatures in natural strunzite, ferristrunzite and ferrostrunzite

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#### Abstract

The temperature variations of the ferric and ferrous hyperfine fields in natural samples of strunzite, ferristrunzite and ferrostrunzite are examined by Mössbauer spectroscopy between 4.2 K and their respective magnetic transition temperatures. The spectra of all three strunzite species are composed of two dominant sextets with equal contribution and with hyperfine fields at 4.2 K of approximately 48.5 T and 48.0 T, respectively. They are assigned to Fe<sup>3+</sup> cations in the Fe(1) and Fe(2) sites in the crystallographic structure. A third magnetic component, with field of 54.8 T ± 0.1 T, is well resolved in the 4.2 K spectrum of ferristrunzite and is attributed to Fe<sup>3+</sup> in the Mn sites of the structure. This third ferric contribution also appears in the spectra of strunzite and ferrostrunzite, however with smaller relative spectral areas as compared to ferristrunzite, i.e., ~10% and ~3% of total spectral area, respectively, where it is 30% for ferristrunzite. For ferrostrunzite, two additional magnetic components are recognized. They are both attributed to Fe<sup>2+</sup> substituting in the Mn sites.

26 Their hyperfine fields at 4.2 K are  $21.8 \text{ T} \pm 0.5 \text{ T}$  and  $28.6 \text{ T} \pm 0.5 \text{ T}$ , respectively, and their  
27 relative contributions 11% and 15%. With increasing temperature, enhanced asymmetric  
28 broadening of the absorption lines is observed for all three samples. This phenomenon is due  
29 to relaxation of the magnetic moments, thus causing gradually broader distribution ranges of  
30 the respective hyperfine fields. The spectra have been fitted with superpositions of model-  
31 independent hyperfine-field distributions. The temperature dependencies of the maximum-  
32 probability hyperfine fields of the two ferric subspectra associated with the Fe(1) and Fe(2)  
33 sites are interpreted in terms of the molecular field theory taking into account the occurrence  
34 of exchange magnetostriction. The variations with temperature of the magnetic hyperfine  
35 fields acting at the Mn sites indicate that the magnetization of the Mn sublattice shows re-  
36 entrant spin glass behaviour. This feature is most prominent for the strunzite sample.

### 37 **1. Introduction**

38 Strunzite, ferristrunzite and ferrostrunzite are three isostructural minerals with triclinic  
39 crystallographic structure (P1) (Fron del 1957, Fanfani et al. 1978, Peacor et al. 1983). The  
40 ideal chemical formulas are  $\text{Mn}^{2+}\text{Fe}^{3+}_2(\text{PO}_4)_2(\text{OH})_2 \cdot 6(\text{H}_2\text{O})$ ,  
41  $\text{Fe}^{3+}\text{Fe}^{3+}_2(\text{PO}_4)_2(\text{OH})_2 \cdot [(\text{H}_2\text{O})_5(\text{OH})]$  (Vochten and De Grave 1990) and  
42  $\text{Fe}^{2+}\text{Fe}^{3+}_2(\text{PO}_4)_2(\text{OH})_2 \cdot 6(\text{H}_2\text{O})$  (Vochten et al. 1995), respectively. Ferristrunzite may be  
43 regarded as the fully oxidized form of ferrostrunzite.

44 The strunzite structure (Figure 1) consists of infinite chains of octahedral ferric sites along the  
45 c axis which are linked to one another by sharing hydroxyl groups and by  $\text{PO}_4$  tetrahedra. The  
46 latter also bind adjacent chains, thus forming slabs that are connected to each other by Mn  
47 octahedra between remaining  $\text{PO}_4$  vertices. Within the chains two ferric sites Fe(1) and Fe(2)  
48 alternate, of which Fe(1) is somewhat more distorted than Fe(2) as indicated by a slight  
49 difference in average Fe-O bond length and average O-Fe-O bond angle (Vochten and De  
50 Grave 1990). The crystallographic unit cell contains two octahedra of each type.

51 In ferristrunzite the manganese is substituted by  $\text{Fe}^{3+}$ , in ferrostrunzite by  $\text{Fe}^{2+}$ . In the first  
52 case the charge balance is re-established by substitution of an  $\text{H}_2\text{O}$  by an  $\text{OH}^-$  at a non-  
53 bridging vertex of the Mn octahedron. In the following, this position is referred to as the Mn  
54 site.

55 Mössbauer spectra (MS) at room temperature (RT) and at 4.2 K for these three mineral  
56 species have been reported earlier (Vochten and De Grave 1990, Vochten et al. 1995, Van  
57 Tassel and De Grave 1992). In this paper the detailed temperature dependencies of the  
58 Mössbauer parameters between 4.2 K and their respective magnetic transition temperatures  
59 ( $T_N$ ) are presented and discussed, with emphasis on the magnetic hyperfine fields, in order to  
60 gain an insight into the relation between the crystallographic structure and the magnetic  
61 properties of these minerals.

62 Preliminary numerical results were reported earlier (Van Alboom and De Grave 2005).  
63 However, the data at that time could not be explained satisfactorily and no sound conclusions  
64 were drawn.

## 65 **2. Experimental**

66 The samples used in this study were described earlier (Vochten and De Grave 1990,  
67 Vochten et al. 1995, Van Tassel and De Grave 1992) and were characterised by these authors  
68 using X-ray diffraction, chemical analysis and Mössbauer spectroscopy at room temperature  
69 (RT) and 4.2 K.

70 In the present work, Mössbauer spectra for powder samples were collected at variable  
71 temperatures in the range where the samples exhibit magnetic ordering. Prior to these  
72 experiments, the respective magnetic order-disorder transition temperatures ( $T_N$ ) were  
73 determined by the Mössbauer thermoscaning technique, in which the temperature ( $T$ )  
74 variation of the  $\gamma$ -ray transmission through the absorber is recorded with the source at zero  
75 velocity (Chambaere and De Grave 1984).

76 For the acquisition of MS conventional time-mode spectrometers were used operating  
77 with a triangular reference signal. The sources were  $^{57}\text{Co}$  in Rh matrix, but all center shift ( $\delta$ )  
78 values quoted hereafter are relative to  $\alpha$ -Fe at room temperature.

79 The obtained spectra were numerically analyzed assuming either discrete Lorentzian-  
80 shaped magnetic components (fitting program referred to hereafter as SFP) or, where found to  
81 be recommended, by model-independent distributions (Vandenberghe et al. 1994) of such  
82 components (program DFP). The positions and intensities of the various absorption lines were  
83 calculated by diagonalization of the full nuclear interaction hamiltonian (Hoy and Chandra  
84 1967, Kündig 1967). Most generally, a magnetically split  $^{57}\text{Fe}$ -MS contains a maximum of  
85 eight absorption lines of which the respective positions are determined by the so-called  
86 Mössbauer parameters, particularly: the center shift  $\delta$ , the hyperfine field  $B_{hf}$  and the polar  
87 ( $\Omega$ ) and the azimuthal angle ( $\Psi$ ) of the hyperfine field in the principal axis system ( $x,y,z$ ) of  
88 the electric field gradient (EFG) with components ( $V_{xx}$ ,  $V_{yy}$ ,  $V_{zz}$ ). Finally, two additional  
89 parameters, related to the quadrupole interaction, affect the line positions, i.e., the quadrupole  
90 splitting  $\Delta E_Q$  and the asymmetry parameter  $\eta$ . They are given by, respectively,

$$91 \quad \Delta E_Q = \frac{1}{2} e^2 q Q = \frac{1}{2} V_{zz} e Q, \quad (1)$$

92 and

$$93 \quad \eta = \frac{V_{xx} - V_{yy}}{V_{zz}}, \quad (2)$$

94 with  $V_{zz}$  the principal EFG component and  $Q$  the nuclear quadrupole moment.

95 However, in the case of  $\text{Fe}^{3+}$  the quadrupole interaction is generally weak in  
96 comparison with the magnetic dipole interaction. As a result, the peak positions are not  
97 independently determined by  $\Delta E_Q$  and  $\Omega$ , but instead by the quadrupole shift ( $2\varepsilon_Q$ ), which is  
98 given by

99 
$$2\varepsilon_Q = \frac{1}{2} \Delta E_Q (3 \cos^2 \Omega - 1). \quad (3)$$

100 Further in this case, the number of lines is reduced to six (sextet) with ideally a mutual line-  
101 intensity ratio of 3:2:1:1:2:3 for a random-powder absorber.

102 In addition to the above mentioned parameters the line width ( $I$ : full width at half  
103 maximum height) and the relative area ( $RA$ ) for each spectral component appearing in a multi-  
104 phase spectrum are determined by fitting the experimental MS.

105 Because of the diversity in types of Mössbauer spectra, four different fitting programs were  
106 used in this study. SFP1 and SFP2 (single component fit programs) were aimed to analyse MS  
107 consisting of a superposition of a number of well-defined discrete magnetically split  
108 subspectra, which are composed of only one subcomponent. Distribution fit programs (DFP1  
109 and DFP2) were applied if subspectra themselves were the superposition of a number of  
110 magnetically split subcomponents related to a distribution of hyperfine fields. In these cases,  
111 only model independent distributions of hyperfine fields (Vandenberghe et al. 1994) were  
112 considered. In SFP1 and DFP1, the positions of the absorption lines (either discrete sextets or  
113 distributions of sextets, respectively) are calculated from their analytical expressions as  
114 function of the various Mössbauer parameters. This approach applies to Fe sites for which the  
115 quadrupole interaction is weak compared to the magnetic hyperfine interaction. If the latter  
116 condition is not met, the programs SFP2 and DFP2 are used which are both based on the  
117 diagonalization of the full nuclear interaction hamiltonian.

### 118 **3. Mössbauer analysis**

#### 119 3.1 Thermoscanning

120 The Mössbauer thermoscanning curves for the three strunzite varieties are reproduced  
121 in Figure 2. Expanding on previously reported data (Van Alboom and De Grave 2005), the  
122 magnetic order-disorder transition temperatures were accurately and consistently determined,  
123 yielding the values  $50.5 \pm 0.5$  K,  $44 \pm 1$  K and  $45 \pm 3$  K for respectively strunzite,

124 ferristrunzite and ferrostrunzite. Hence, the Fe-for-Mn substitution in the strunzite structure  
125 seems to weaken the strength of the overall magnetic exchange interaction. The  $T_N$ -value as  
126 determined from thermoscaning gives an overall value for a sample, and therefore does not  
127 correspond necessarily to the values of all the magnetic sublattices of the sample.

### 128 3.2 Ferristrunzite

129 At 4.2 K the experimental MS of ferristrunzite was satisfactorily described using SFP1  
130 assuming a superposition of three ferric Lorentzian shaped sextets, S1, S2, and S3 (Vochten  
131 and De Grave 1990), attributed to the three distinct Fe sites available in the ferristrunzite  
132 structure. The respective center shift values are very similar ( $\delta \approx 0.50$  mm/s) and are  
133 consistent with an octahedral co-ordination.

134 Two of these sextets (S1 and S2) have slightly different hyperfine fields ( $B_{hf} = 48.4$  T  
135  $\pm 0.1$  T and  $B_{hf} = 47.8$  T  $\pm 0.1$  T) while the third one (S3) has a remarkably higher hyperfine  
136 field ( $B_{hf} = 54.8$  T  $\pm 0.1$  T) at 4.2 K. S1 and S2 are assigned to  $\text{Fe}^{3+}$  at the Fe(1 or 2) and Fe(2  
137 or 1) sites in the structure. Although the quadrupole shifts for  $\text{Fe}^{3+}$ (S1) and  $\text{Fe}^{3+}$ (S2) are  
138 different, further unambiguous assignment to the Fe(1) or Fe(2) sites isn't possible, because  
139 the difference in  $2\varepsilon_Q$  is related to the difference in orientation of the local hyperfine field with  
140 respect to the local EFG in both sites, of which the orientation isn't known until now. S3 is  
141 attributed to  $\text{Fe}^{3+}$  at the Mn position. The center shifts corresponding to S1 and S2 were  
142 constrained to have equal values. The subspectral area fractions were fixed on the ratio  
143 0.35:0.35:0.30 which corresponds to a model for which the aluminum, which was detected by  
144 EDX to be present in the sample by an amount of 0.10 per formula unit, is located on the Mn  
145 sites in the structure. It was experienced, however, that a forced random distribution of the Al  
146 among the three available lattice sites or among sites Fe(1) and Fe(2) did not significantly  
147 affect the adjusted hyperfine parameter values. The line area ratios 3:x:y of the outer to  
148 middle to inner lines for a particular subspectrum were assumed to be the same for all three

149 components and were found to be 3:2.43:1.14. The deviation from the ideal 3:2:1 ratio can be  
150 due to thickness effects or texture effects, however here, also relaxation effects cannot be  
151 excluded (see further). As it is experienced that the deviation is irrelevant for the subsequent  
152 determination and interpretation of the temperature variations of the various hyperfine  
153 parameters, the hyperfine field in particular, no further attention is paid to the line area ratios.

154 At 8 K and higher temperatures  $T$ , significant line broadening is noticed, as well as  
155 additional absorption at low velocities which becomes more prominent with increasing  $T$ . It  
156 was found that its contribution could be taken into account by introducing an extra doublet  
157 component D1 in the foregoing analyzing procedure, hereby keeping the same constraints for  
158 the sextets as those imposed on the fit of the 4.2 K spectrum. The adjusted values of the D1  
159 center shift values are, within experimental error limits, equal to those of the sextet  
160 components. Further, the MS acquired at 8 K and 14 K show strongly overlapping sextet  
161 subspectra S1, S2, and S3 and consequently their Mössbauer parameters could not be  
162 determined accurately using the SFP1 fitting approach. Nevertheless, a remarkably different  
163  $T$ -dependence of the hyperfine fields among these subspectra is observed. While  $B_{\text{hf},S1}$  and  
164  $B_{\text{hf},S2}$  show nearly the same  $T$ -dependence,  $B_{\text{hf},S3}$  decreases more rapidly with increasing  $T$   
165 than  $B_{\text{hf},S1}$  and  $B_{\text{hf},S2}$  and above 14 K the field  $B_{\text{hf},S3}$  becomes smaller than the corresponding  
166 S1 and S2 values.

167 From 17 K on, the asymmetric broadening of the absorption lines is considerable and  
168 increases drastically with increasing  $T$ . As a result, the analysis of these MS using a  
169 superposition of three discrete sextets and one doublet (SFP1 approach) failed. Therefore, the  
170 theoretical spectra were calculated by the superposition of three model independent  
171 distributions of hyperfine fields, respectively, related to the already mentioned Fe sites, using  
172 the distribution fitting program DFP1. In this case, a subspectrum is generally the  
173 superposition of a certain number of subcomponents (given by #div), of which the hyperfine

174 fields have different values as given by the distribution, whereas the center shifts, the  
175 quadrupole shifts, the line widths and the corresponding line area ratios are constraint to equal  
176 values for all subcomponents. Here the line area ratios of all these subcomponents were fixed  
177 at 3:2:1. The central component D1 was fitted as a discrete, Lorentzian-shaped quadrupole  
178 doublet. At 17 K and higher  $T$ , the center shifts of all subspectra were constrained to have  
179 equal values. Regrettably, in the minimization routine used in the distribution program for  
180 fitting the MS, it is not possible to impose fixed values for the  $RA$  of the subspectra. Such  
181 possibility would indeed be very useful when strongly overlapping subspectra determine the  
182 overall Mössbauer line shape. As a consequence and also because of the nearly total overlap  
183 of S1 and S2, the  $RA$  results for S1 and S2 fluctuated strongly depending on the input  
184 parameter values. Therefore, at each  $T$ , several trial fits were done with different distribution  
185 ranges (lower and upper limits for the allowed hyperfine fields) and/or different fixed  $2\epsilon_Q$   
186 values for S1 and S2 in order to get nearly equal  $RA$  values for S1 and S2. It was found that  
187 fixing or not fixing the elemental line width  $\Gamma$ , only slightly affected the adjusted  $RA$  values,  
188 while the positions of the respective maximum probability hyperfine fields did not change  
189 substantially.

190         Applying the calculation approaches as outlined above, the MS at temperatures  
191 between 4.2 K and  $T_N$  could satisfactorily be reproduced. Figure 3 shows selected  
192 experimental and calculated spectra as well as the probability distribution profiles  $p(B_{hf})$  of  
193 the three sextet components. It is noticed that the  $p(B_{hf})$  profiles for S1 and S2 are very close  
194 to one another, both in shape and in probability for a given field value. In spite of the strong  
195 overlap of the sextet components at higher  $T$ , and hence their ill-defined parameter values,  
196 reasonably consistent results for these parameters were obtained, especially for the hyperfine  
197 fields  $B_{max}$  at maximum probability.

198 Table I gives the results of the relevant Mössbauer parameters for which the  $RA$  of the  
199 S1 and S2 components are in good approximation the same and for which the Mössbauer  
200 parameters of all the subspectra show a consistent and physical acceptable dependence of  
201 temperature. The quadrupole shifts of the various sextets are not sharply defined, but it is  
202 clear that their values for S1 and S2 have opposite signs. Figure 3 contains the experimental  
203 and calculated MS. At 17 K and higher  $T$ , the  $RA$  of S1 and S2 decrease substantially.

204 The quadrupole splitting of component D1 was determined for all  $T$  to be  
205 approximately 0.75mm/s, which is in good agreement with the value of the quadrupole  
206 splitting of the doublet absorption at maximum probability above  $T_N$  (Vochten and De Grave  
207 1990).  $2\varepsilon_{Q,S1}$  and  $2\varepsilon_{Q,S2}$  had always a different sign, and vary between  $-0.63$  and  $-0.42$  mm/s  
208 for S1 and between 0.11 and 0.32 mm/s for S2. Most probably these fluctuations are a  
209 consequence of the overlap of S1 and S2, by which it was not possible to determine these  
210 values more precisely. A change in sign of  $2\varepsilon_{Q,S3}$  below 8 K and above 14 K is determined.

211 Because the sign of  $2\varepsilon_Q$  is related to the orientation of the local hyperfine field to the principal  
212 axis of the EFG, the difference in sign of  $2\varepsilon_{Q,S1}$  and  $2\varepsilon_{Q,S2}$  indicates a difference in that  
213 orientation at Fe(1) and Fe(2) sites. For the same reason the difference in sign of  $2\varepsilon_{Q,S3}$  below  
214 8 K and above 14 K should be interpreted as a change in the direction of the local hyperfine  
215 field with respect to the principal axis of the local EFG. However, taking into account the  
216 constraints on some of the parameters during the analysis, one best takes any reservation  
217 concerning this finding.

218 The effect of increasing  $RA_{D1}$  and decreasing  $RA$  of the other components with increasing  $T$   
219 above 17 K, is interpreted in the context of relaxation of the magnetic moments, by which the  
220 magnetic moments remain increasingly less time in their equilibrium positions as  $T$  increases.  
221 When the relaxation time is sufficiently shorter than the nuclear Larmor precession time (of  
222 the order  $10^{-7}$  s), then the Mössbauer absorption is of doublet type. In the opposite case, when

223 the magnetic moments stay long enough in their equilibrium state, one observes a  
224 magnetically split spectrum. In the intermediate case, the absorption manifests itself as the  
225 superposition of a continuum of hyperfine fields between 0 T (doublet situation) and a  
226 maximum value.

227 From these considerations, the D1 component has to be associated with  $\text{Fe}^{3+}$  for which the  
228 relaxation time is already short enough to get a more doublet-like absorption. That the  
229 relaxation phenomenon is spread over a relatively wide temperature range, means that the  
230 magnetic exchange interaction with neighbouring magnetic ions, which is responsible for the  
231 orientation of the magnetic moments, is not equally strong at each site, and being weakest for  
232  $\text{Fe}^{3+}$  at Mn sites. This last assertion can be easily understood because the Mn site in strunzite  
233 is a relatively isolated octahedrally co-ordinated site between the slabs of chains of octahedral  
234 ferric sites in the crystallographic structure. Therefore it is plausible that the  $\text{Fe}^{3+}$  at the Mn  
235 position has weaker magnetic exchange interactions with neighbouring magnetic ions than  
236  $\text{Fe}^{3+}(1)$  and  $\text{Fe}^{3+}(2)$  do.

237 The broadening of the absorption lines is interpreted as being the consequence of local  
238 differences in the strength of the magnetic exchange interactions between neighbouring  
239 magnetic ions caused by local differences in the co-ordination of the probe ions. Consequently  
240 the  $T$ -dependence of the hyperfine field at a certain position can be different from the one at  
241 another position of the same site, whereas the saturation values at lowest  $T$  are approximately  
242 the same. In the former case, a broad  $B_{hf}$  distribution range arises and an asymmetric  
243 broadening of the outer absorption lines of the spectral component is observed when  $T$   
244 increases. In order to analyse spectra showing relaxation effects, a special fitting routine is  
245 designed in the past (Blume and Tjon 1968), which is successfully applied to Mössbauer  
246 spectra with single component subspectra. However because of the strong overlap of the

247 subspectra, which themselves show a distribution of hyperfine fields, this fitting routine is not  
248 applicable for the minerals in this study.

### 249 3.3 Strunzite

250 Using SFP1 the MS of strunzite at 4.2 K was calculated by three sextets and an additional  
251 doublet component D2 at small velocities with  $RA$  less than 3 %. This doublet was assigned to  
252 some iron in the window of the equipment. The center shifts and line widths of S1 and S2  
253 were constrained to equal values. Broadening of the outer lines of these components was  
254 taken into account by a linear broadening parameter ( $\Delta\Gamma$ ): the inner line to middle to outer  
255 lines having line widths  $\Gamma$ ,  $\Gamma+\Delta\Gamma$  and  $\Gamma+2\Delta\Gamma$ . Additionally a coupled 3:x:1 ratio for the line  
256 area of the sextet absorption lines was used. In this way the MS was satisfactorily reproduced  
257 and the results for the Mössbauer parameters of the sextets at 4.2 K (see Table II) were very  
258 similar to those at 4.2 K in ferristrunzite. In comparison with ferristrunzite and the  
259 components S1 and S2, the component with the highest  $B_{hf}$  (S3) had now a substantially  
260 smaller area fraction of 0.11 approximately, S1 and S2 having equal area fractions of 0.44  
261 approximately. However the sign of  $2\varepsilon_{Q,S3}$  at 4.2 K in strunzite was opposite to the  
262 corresponding one in ferristrunzite.

263 The MS at 10 K was also satisfactorily reproduced by an analogous superposition of discrete  
264 components for which the  $RA$  of the sextets were fixed at 0.44:0.44:0.11 and the Mössbauer  
265 parameters of D2 at the 4.2 K values (except the line width). In this case S3 overlapped totally  
266 with S1 and S2, and its parameters therefore are somewhat more inaccurate.

267 At increasing temperatures,  $B_{hf,S3}$  decreased very rapidly and already at 15 K, its value was  
268 much smaller than  $B_{hf,S1}$  and  $B_{hf,S2}$ . Meanwhile, using SFP1, the line width of S3 increased  
269 strongly with increasing  $T$ , most probably because of relaxation effects. Therefore, from 15 K  
270 and higher  $T$ , the MS were analysed using DFP1 with the distribution fitting model for S3,  
271 whereas S1 and S2 were reproduced by discrete sextet components with a coupled linear

272 broadening of the absorption lines. All sextet components were constrained to equal 3:x:1 line  
273 area ratios and to equal center shifts. However DFP1 has the disadvantage that the relative  
274 areas of the components could not be fixed. Up to 20 K the parameters of D2 were fixed to  
275 the values found at 4.2 K. Above 20 K up to 47 K the lower limit of the distribution range for  
276  $B_{hf,S3}$  was lowered to 0 and consequently overlapped the absorption corresponding to D2. A  
277 separate D2 doublet therefore was not taken into consideration for these  $T$ . At 43 and 47 K the  
278 MS showed also asymmetric broadening for the S1 and S2 absorption. Therefore these MS  
279 were reproduced using DFP1 with distribution models for the hyperfine fields related to the  
280 S1, S2 and S3 absorption. At 47 K additional doublet absorption D1 was resolved. As for  
281 ferristrunzite, at each  $T$ , several fits were done with different distribution ranges and/or  
282 different fixed  $2\epsilon_Q$  values for S1 and S2 in order to get nearly equal RA for S1 and S2. Again  
283 it was clear that the dominating sextet absorption was the superposition of two components  
284 with very similar hyperfine field distributions and with a different sign for the related  
285 quadrupole shifts, indicating that the hyperfine field orientation is different with respect to the  
286 local principal axis of the EFG at the corresponding sites. At 43 K and at 47 K the subspectra  
287 in the MS are clearly collapsing and at these  $T$ , the highest value of the hyperfine field  
288 distribution of S3 was only 12.0 T and 10.0 T, respectively. The results for which the RA of  
289 the S1 and S2 components are in good approximation the same and for which the Mössbauer  
290 parameters of all the subspectra show a consistent and physical acceptable dependence of  
291 temperature, are given in Table II. Figure 4 contains the experimental and calculated MS.

### 292 3.4 Ferrostrunzite

293 In the MS of ferrostrunzite below  $T_N = 45 \pm 3$  K, because of the presence of  $\text{Fe}^{2+}$  in the  
294 sample, magnetically split ferrous subspectra were observed in addition to the  $\text{Fe}^{3+}$   
295 subspectra. Because for  $\text{Fe}^{2+}$ , the quadrupole interaction and the magnetic dipole interaction  
296 are of the same order, the MS of ferrostrunzite below  $T_N$  were analysed with the method based

297 on the diagonalization of the full nuclear interaction hamiltonian. At 5 K the MS could be  
298 reproduced by three discrete ferric (sextet) and two ferrous magnetically split subspectra  
299 using SFP2. A position dependent width broadening parameter ( $\Delta I$ ) accounts for some  
300 broadening of the outer lines caused by small local variations of  $B_{hf}$ . As for the other  
301 samples, two of these ferric subspectra (S1 and S2) are related to  $Fe^{3+}$  at the Fe(1 or 2) and  
302 Fe(2 or 1) sites and the third ferric component (S3) is assigned to ferric ions at the Mn site.  
303 Keeping the area fractions of S1 and S2 equal to each other during the fitting process gives  
304 no worse goodness-of-fit values, thus makes an equal ferric occupation of both sites very  
305 plausible. Because the ferrous quadrupole splitting is temperature dependent as consequence  
306 of the varying Boltzmann occupation of the lower  $^5D$  orbital electron energy levels, its value  
307 saturates at low temperatures; this occurs already at about 80 K. From this consideration,  
308 quadrupole splittings of the two ferrous magnetically split subspectra (S4 and S5) were fixed  
309 at the values just above the magnetic transition temperature, 3.04 and 2.52 mm/s respectively  
310 (Van Tassel and De Grave 1992). The area fractions of both ferrous subspectra were fixed at  
311 the ratio determined just above the magnetic transition temperature. This was also done for  
312 the center shifts of both components.

313 At 22 K and 28 K an asymmetric broadening of the outer absorption peaks was visually  
314 observed, indicating an inhomogeneous hyperfine field decrease at Fe(1) and Fe(2) sites.  
315 Therefore, at these temperatures, the MS were analysed using ferric subspectra with  
316 distributions of hyperfine fields. Also an increase of the absorption near zero velocity was  
317 observed and S3 could not be resolved anymore. In order to reproduce the MS at these  $T$ ,  
318 therefore a magnetically split subspectrum (S6) was added, for which the hyperfine field  
319 range extends from 0.0 to 10.0 T. The quadrupole shift of it was fixed at the value of 0.755  
320 mm/s for the quadrupole splitting with highest probability in the paramagnetic spectrum just  
321 above  $T_N$  (Van Tassel and De Grave 1992). Because of the strong overlap of the subspectra,

322 by which their resolution is strongly limited, the S6 absorption is composed of the collapsing  
323 S3 on the one hand, and parts of the S1 and S2 absorption on the other hand. Center shifts and  
324 widths of the ferric subspectra were constrained to equal values. Because of the strong  
325 overlap with the other subspectra, the ferrous absorption was reproduced by single  
326 magnetically split subspectra with large line widths constrained equal to each other. For the  
327 calculations at these  $T$ , DFP2 was used. However relative areas could not be fixed, and  
328 because of the strong overlap with the ferric subspectra, the ferrous ones were not well  
329 resolvable at these  $T$ . Therefore, the parameters determined for them are less accurate. The  
330 results of all these calculations for which the  $RA$  of the S1 and S2 components are in good  
331 approximation the same and for which the Mössbauer parameters of all the subspectra show a  
332 consistent and physical acceptable dependence of temperature, are given in Table III. Figure 5  
333 contains the experimental and calculated MS. It is clear that the area fractions of S1 and S2  
334 decrease above 22 K.

#### 335 **4. Hyperfine field analysis**

##### 336 *4.1 Temperature dependence of the hyperfine field*

337 As is generally known, the  $T$ -dependence of the sublattice magnetisation ( $\sigma$ ) in a classical  
338 magnetic structure can be described by a Brillouin curve within the molecular field theory of  
339 magnetism (Morrish 1965):

$$340 \quad \sigma = B_{\frac{5}{2}}(\chi) = \frac{6}{5} \coth\left(\frac{6}{5}\chi\right) - \frac{1}{5} \coth\left(\frac{1}{5}\chi\right) \quad \text{with} \quad \chi = \frac{15 T_N}{7 T} \sigma$$

341 Because the hyperfine field  $B_{hf}$  is in very good approximation proportional to the sublattice  
342 magnetization, here related to an iron site, the formula can also be applied to describe the  $T$ -  
343 dependence of  $B_{hf}$ :

$$344 \quad \sigma = \frac{B_{hf}(T)}{B_{hf}(0K)} \Leftrightarrow B_{hf}(T) = B_{hf}(0K) \times B_{\frac{5}{2}}(\chi)$$

345 These  $B_{hf}(T)$  Brillouin curves are determined by the mean spin value of the probe-ions (2.5 for  
346 ferric ions), the saturation value of the magnetization at low temperatures (here: the hyperfine

347 field at lowest  $T$ ) and the corresponding magnetic transition temperature (here: the  $T_N$  value as  
348 determined from thermoscanning). Figure 6 shows the experimental  $T$ -dependences of  $B_{hf}$   
349 (most probable values) and the calculated Brillouin curves for the respective ferric sites and  
350 samples. It is clear that the molecular field model gives mostly only a raw description of the  
351  $T$ -dependences of  $B_{hf,S1}$  and  $B_{hf,S2}$ , and is totally deficient for  $B_{hf,S3}(T)$ . However for strunzite  
352 and ferrostrunzite a better description of  $B_{hf,S1}$  and  $B_{hf,S2}$  as function of  $T$  was obtained by a  
353 curve which takes into account the exchange magnetostriction associated with a non-second-  
354 order magnetic transition (Bean and Rodbell 1962). The argument of the Brillouin curve is  
355 than given by:

$$356 \quad \chi = B_{\frac{S}{2}}^{-1}(\sigma) = [-2a - 4b\eta_{nso}\sigma^2] \frac{T_N}{T} \sigma$$
$$a = -\frac{15}{14} \quad b = -\frac{2331}{9604}$$

357 Such a curve is calculated using an extra parameter, here noted by  $\eta_{nso}$  in order to distinguish  
358 it from the asymmetry parameter.  $\eta_{nso}$  is a measure for the departure from an ideal second-  
359 order magnetic transition and is determined by the fractional volume change at the transition,  
360 the compressibility, and the slope of the variation of the transition temperature with the lattice  
361 volume (Bean and Rodbell 1962). In the case of an ideal second-order magnetic transition  $\eta_{nso}$   
362 = 0 and the curve is then the classical Brillouin curve;  $\eta_{nso} > 1$  corresponds to a first-order  
363 transition.  $\eta_{nso}$  can be determined by fitting this kind of curve to the experimental data using  
364 the same values for the other parameters as for the foregoing calculation of the corresponding  
365 classical Brillouin curve. Because the  $T$ -dependences of  $B_{hf,S1}$  and  $B_{hf,S2}$  are mutually very  
366 similar in strunzite, one finds  $\eta_{nso} = 0.43 \pm 0.01$  for both S1 and S2 in the case of strunzite, and  
367  $\eta_{nso,S1} = 0.37 \pm 0.02$  and  $\eta_{nso,S2} = 0.34 \pm 0.02$  in the case of ferrostrunzite, indicating that the  
368 magnetic transition at the corresponding  $T_N$  is not a pure second-order. For ferristrunzite the  
369 best agreement with the experimental data was obtained for  $\eta_{nso} = 0 \pm 0.02$ , yielding the

370 classical Brillouin curve. Because the fractional volume changes at the magnetic transition  
371 temperatures for strunzite and ferrostrunzite are not known, more quantitative data about  
372 compressibility and slope of the variation of the transition temperature with the lattice volume  
373 cannot be derived from the obtained values for  $\eta_{nso}$ . In ferrostrunzite (Figure 3) it can be seen  
374 that  $T_N(S1,S2)$  is different from  $T_N(S3,S6)$ , demonstrating that different sublattices can have  
375 different magnetic transition temperatures.

#### 376 *4.2 Re-entrant spin glass behaviour*

377 The  $B_{hf,S3}$   $T$ -dependence is remarkable. In strunzite (Figure 6)  $B_{hf,S3}$  (most probable values)  
378 decreases strongly when  $T$  increases from 0 to about 20 K, after which  $B_{hf,S3}$  decreases less  
379 quickly when  $T$  further increases up to  $T_N$ . This kind of  $T$ -dependence is typical for re-entrant  
380 spin glass behaviour (De Bakker et al. 1997). This phenomenon occurs in systems in which  
381 some spins, here the ferric ones at the Mn sites, are frustrated in the sense that they are liable  
382 to conflicting magnetic exchange interactions which never can be satisfied simultaneously for  
383 all interactions (Stein 1989). The kink in the curve determines the so-called spin glass freezing  
384 temperature ( $T_{sg}$ ). In strunzite,  $T_{sg}$  is estimated at  $21 \pm 1$  K. Above  $T_{sg}$ , the frustrated spins are  
385 in a local paramagnetic-like state, and  $B_{hf,S3}(T)$  then reflects the normal course of the  
386 spontaneous magnetisation (of non-frustrated  $Fe^{3+}$  spins at the Mn site), which almost reaches  
387 its saturation value at  $T_{sg}$ . Below  $T_{sg}$ , the frustrated spins are frozen in and their contribution is  
388 superimposed on the mean spin value of the non-frustrated spins. Because the hyperfine field  
389 of the non-frustrated spins was approximately saturated at  $T_{sg}$ ,  $B_{hf,S3}(T)$  then mainly reflects  
390 the  $T$ -dependence of the frustrated spins (Huck and Hesse 1989). A less pronounced  
391 analogous tendency is determined for  $B_{hf,S3}$  in ferristrunzite and ferrostrunzite.  
392 More quantitative discussion about this phenomenon can only be done when more data  
393 concerning the  $T$ -dependence of the ferric hyperfine fields are determined and the magnetic

394 structure is better clarified. Complete strunzite-ferristrunzite and strunzite-ferrostrunzite  
395 solution series are interesting mineral series for further research on this subject.

## 396 **5. Conclusion**

397 Strunzite, ferristrunzite and ferrostrunzite exhibit Mössbauer spectra at temperatures below  
398 the magnetic ordering temperature that are rather complicated to analyze. The  $B_{hf}$  of the iron  
399 located at the Mn sites decreases more strongly with temperature than these of the Fe(1) and  
400 Fe(2) sites. When temperature increases sufficiently, the absorption lines of the dominant  
401 sextets become asymmetrical and broaden as a consequence of relaxation effects. In strunzite  
402 this becomes visible above 40 K, in ferristrunzite and ferrostrunzite already above about 17 K  
403 and 23 K respectively. The  $T$ -dependence of the hyperfine fields of the  $Fe^{3+}(1)$  and  $Fe^{3+}(2)$   
404 sites in strunzite and ferrostrunzite is successfully described by a model which takes exchange  
405 magnetostriction effects into account and shows that the magnetic transition is not of pure  
406 second order. However in ferristrunzite, a classical Brillouin dependence is more likely. The  
407  $T$ -dependence of the hyperfine field of the  $Fe^{3+}$  at the Mn sites in strunzite is rather  
408 extraordinary and shows re-entrant spin glass behaviour.

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412

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- 451

452 **Figure captions**

453 Figure 1. Part of the crystallographic structure of strunzite viewed along the *c* axis (adapted  
454 from Fanfani 1978).

455 Figure 2. Mössbauer thermoscanning curves of the three strunzite varieties.

456 Figure 3a. Experimental and calculated MS for ferristrunzite at 4.2 K. S1: blue; S2: purple;  
457 S3: magenta; calculated MS: red. Parameters are in Table I.

458 Figure 3b. Experimental and calculated MS for ferristrunzite (left) and corresponding  
459 probability histograms for the hyperfine fields (right) at selected temperatures. S1: blue; S2:  
460 purple; S3: magenta; D1: orange; calculated MS: red. Parameters are in Table I.

461 Figure 4. Experimental and calculated MS for strunzite at selected temperatures. S1: blue; S2:  
462 purple; S3: magenta; D1: orange; D2: pink; calculated MS: red. Parameters are in Table II.

463 Figure 5a. Experimental and calculated MS for ferrostrunzite at 5 K and 10 K. S1: blue; S2:  
464 purple; S3: magenta; S4: orange; S5: green; calculated MS: red. Parameters are in Table III.

465 Figure 5b. Experimental and calculated MS for ferrostrunzite (left) and corresponding  
466 probability histograms for the hyperfine fields (right) at 22 K and at 28 K. S1: blue; S2:  
467 purple; S4: orange; S5: green; S6: pink; calculated MS: red. Values of the relevant Mössbauer  
468 parameters are listed in Table III.

469 Figure 6. *T*-dependence of the hyperfine fields (most probable values) in ferristrunzite,  
470 strunzite and ferrostrunzite. Dotted lines are calculated Brillouin curves. Full lines are  
471 calculated accounting for exchange magnetostriction (see text) (Bean and Rodbell 1962). S1:  
472 blue; S2: purple; S3: magenta; S6: pink.

473

474 **TABLES**

475 Table I. Mössbauer parameters for ferristrunzite below  $T_N$  at selected temperatures. See text for the meaning of  
 476 the symbols. Except at  $T=4.2$  K (errors in the last digit), estimated errors are: 0.5 K on  $T$ , 0.01 mm/s on  $\delta$ , 0.05  
 477 mm/s on  $2\varepsilon_Q$ , 0.03 mm/s on  $\Delta E_Q$ , 0.02 mm/s on  $\Gamma$ , 0.5 T on  $B_{max}$  ( $B_{hf}$  with maximum probability), 0.5% on RA.  
 478 At 4.2 K the given widths refer to the outer lines of the sextets. SFP1 (single component fit program), DFP1  
 479 (distribution fit program) are both based on analytical expressions for the calculation of the absorption line  
 480 positions.  $B_{hfmin}$ ,  $B_{hfmax}$ , #div: underlimit, upperlimit and number of divisions of the distribution range of  $B_{hf}$  for  
 481 use in DFP1.

482

$T$ (K)	Progr	Comp	$B_{hfmin}$ (T)	$B_{hfmax}$ (T)	#div	$B_{max}$ (T)	$\delta$ (mm/s)	$2\varepsilon_Q$ (mm/s)	$\Delta E_Q$ (mm/s)	$\Gamma$ (mm/s)	RA (%)
4.2	SFP1	S1				48.4(1)	0.508(5)	-0.63(2)		0.48(1)	35
		S2				47.8(1)	0.508(5)	0.11(2)		0.42(1)	35
		S3				54.8(1)	0.498(5)	-0.22(2)		0.35(1)	30
17	DFP1	S1	35.0	45.5	23	41.7	0.51	-0.48		0.40	24
		S2	35.0	45.5	23	41.8	0.51	0.18		0.40	24
		S3	10.0	35.0	26	31.2	0.51	0.14		0.40	38
		D1					0.51		0.71	0.66	14
21	DFP1	S1	30.0	45.0	31	40.1	0.51	-0.46		0.37	22.5
		S2	30.0	45.0	31	40.4	0.51	0.16		0.37	22.5
		S3	10.0	35.0	26	24.4	0.51	0.18		0.37	37
		D1					0.51		0.74	0.58	18
29	DFP1	S1	20.0	40.0	21	33.8	0.51	-0.46		0.35	16
		S2	20.0	40.0	21	34.7	0.51	0.32		0.35	16
		S3	3.0	31.0	29	14.3	0.51	0.12		0.35	32
		D1					0.51		0.75	0.53	36
33	DFP1	S1	20.0	40.0	21	30.9	0.52	-0.42		0.31	14.5
		S2	20.0	40.0	21	31.6	0.52	0.32		0.31	14.5
		S3	3.0	31.0	29	12.3	0.52	0.17		0.31	29
		D1					0.52		0.76	0.39	42

483

484

485 Table II. Mössbauer parameters for strunzite below  $T_N$ . See text for the meaning of the symbols. Typical errors  
 486 are: 0.5 K on T, 0.005 mm/s on  $\delta$ , 0.005 mm/s on  $\Delta E_Q$  and  $2\varepsilon_Q$ , 0.005 mm/s on  $\Gamma$ , 0.5 T on  $B_{max}$  ( $B_{hf}$  with  
 487 maximum probability), 0.1 on x, 0.5% on RA. SFP1 (single component fit program), DFP1 (distribution fit  
 488 program) are both based on analytical expressions for the calculation of the absorption line positions.  $B_{hfmin}$ ,  
 489  $B_{hfmax}$ , #div: underlimit, upperlimit and number of divisions of the distribution range of  $B_{hf}$  for use in DFP1. For  
 490 the discrete sextets, given line widths are for the outer lines. x is the line area ratio of the second (and fifth) line  
 491 relative to the third (and fourth) line in a sextet. The line area ratio of the first (and sixth) line to the third (and  
 492 the fourth) line of a sextet was always fixed at 3.  
 493

T (K)	Progr	Comp	$B_{hfmin}$ (T)	$B_{hfmax}$ (T)	#div	$B_{max}$ (T)	$\delta$ (mm/s)	$2\varepsilon_Q$ (mm/s)	$\Delta E_Q$ (mm/s)	$\Gamma$ (mm/s)	x	RA (%)
4.2	SFP1	S1				48.6	0.518	-0.425		0.436	2.2	43.3
		S2				48.3	0.518	0.078		0.436	2.2	43.3
		S3				54.2	0.501	0.177		0.490	2.2	10.8
		D2					0.316		0.330	0.802		2.6
10	SFP1	S1				47.2	0.519	-0.354		0.393	2.1	43.5
		S2				47.0	0.519	0.126		0.393	2.1	43.5
		S3				44.1	0.519	0.068		0.687	2.1	10.9
		D2					0.316		0.330	0.540		2.1
15	DFP1	S1				45.8	0.515	-0.240		0.381	2.0	42.8
		S2				45.8	0.515	0.185		0.381	2.0	42.8
		S3	30.0	45.0	31	37.5	0.515	0.413		0.236	2.0	12.0
		D2					0.316		0.330	0.510		2.4
20	DFP1	S1				44.6	0.507	-0.115		0.386	1.7	40.9
		S2				44.6	0.507	0.190		0.386	1.7	40.9
		S3	5.0	35.0	61	22.0	0.507	0.371		0.449	1.7	16.6
		D2					0.316		0.257	0.500		1.6
25	DFP1	S1				43.3	0.510	-0.112		0.384	1.7	41.2
		S2				43.1	0.510	0.197		0.384	1.7	41.2
		S3	0.0	25.0	51	16.7	0.510	0.449		0.250	1.7	17.6
30	DFP1	S1				41.6	0.510	-0.107		0.376	1.7	40.9
		S2				41.5	0.510	0.190		0.376	1.7	40.9
		S3	0.0	25.0	51	13.5	0.510	0.446		0.239	1.7	18.2
35	DFP1	S1				39.1	0.506	-0.100		0.381	1.7	40.0
		S2				38.9	0.506	0.206		0.381	1.7	40.0
		S3	0.0	25.0	51	11.4	0.506	0.376		0.256	1.7	20.0
40	DFP1	S1				35.3	0.508	-0.110		0.450	1.7	40.8
		S2				35.2	0.508	0.190		0.450	1.7	40.8
		S3	0.0	20.0	41	9.2	0.508	0.643		0.260	1.7	18.4
43	DFP1	S1	25.0	35.0	21	31.8	0.508	-0.117		0.231	1.8	39.8
		S2	25.0	35.0	21	31.5	0.508	0.223		0.231	1.8	39.8
		S3	0.0	12.0	25	6.9	0.508	0.550		0.231	1.8	20.4
47	DFP1	S1	15.0	28.0	14	24.5	0.518	-0.100		0.286	1.8	35.1
		S2	15.0	28.0	14	24.2	0.518	0.229		0.286	1.8	35.1
		S3	0.0	10.0	21	5.1	0.518	0.452		0.286	1.8	19.5
		D1					0.518		0.749	0.396		10.3

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496 Table III. Mössbauer parameters for strunzite below  $T_N$ . See text for the meaning of the symbols. Typical errors  
 497 are: 0.5 K on  $T$ , 0.01 mm/s on  $\delta$ , 0.01 mm/s on  $\Delta E_Q$  and  $2\varepsilon_Q$ , 0.10 on  $\eta$ , 0.01 mm/s on  $\Gamma$ ,  $5^\circ$  on  $\Omega$ , 0.5 T on  $B_{max}$   
 498 ( $B_{hf}$  with maximum probability), 0.5% on RA. SFP2 (single component fit program) and DFP2 (distribution fit  
 499 program) are both based on the diagonalization of the full nuclear interaction hamiltonian.  $B_{hfmin}$ ,  $B_{hfmax}$ , #div:  
 500 underlimit, upperlimit and number of divisions of the distribution range of  $B_{hf}$  for use in DFP2.  
 501

$T$ (K)	Progr	Comp	$B_{hfmin}$ (T)	$B_{hfmax}$ (T)	#div	$B_{max}$ (T)	$\delta$ (mm/s)	$2\varepsilon_Q$ (mm/s)	$\Delta E_Q$ (mm/s)	$\eta$	$\Omega$ ( $^\circ$ )	$\Gamma$ (mm/s)	RA (%)
5	SFP2	S1				48.2	0.51	-0.30				0.54	35.8
		S2				47.9	0.51	0.10				0.54	35.8
		S3				55.3	0.51	0.16				0.30	2.7
		S4				21.8	1.37		3.04	0.30	80	1.02	10.9
		S5				28.6	1.37		2.52	0.30	80	1.02	14.8
10	SFP2	S1				47.4	0.51	-0.34				0.56	35.5
		S2				47.2	0.51	0.04				0.56	35.5
		S3				46.7	0.51	0.30				0.25	2.7
		S4				16.8	1.37		3.04	0.08	68	0.90	11.1
		S5				22.7	1.37		2.52	0.08	80	0.90	15.2
22	DFP2	S1	35.0	45.5	21	44.4	0.52	-0.17				0.39	33.3
		S2	35.0	45.5	21	44.0	0.52	0.06				0.39	33.3
		S4				8.7	1.37		3.04	0.09	76	0.59	12.1
		S5				9.0	1.37		2.52	0.09	85	0.59	14.2
		S6	0.0	10.0	20	7.1	0.52	0.76				0.39	7.1
		S1	30.0	43.0	26	41.4	0.51	-0.19				0.33	31.8
28	DFP2	S2	30.0	43.0	26	41.0	0.51	0.16				0.33	31.8
		S4				6.5	1.37		3.04	0.01	79	0.44	10.2
		S5				5.9	1.37		2.52	0.01	80	0.44	16.5
		S6	0.0	10.0	20	3.1	0.51	0.76				0.33	9.7

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**REVISION 1**  
**FIGURES**

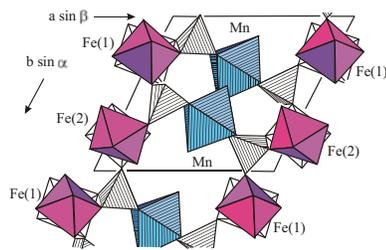


Figure 1.

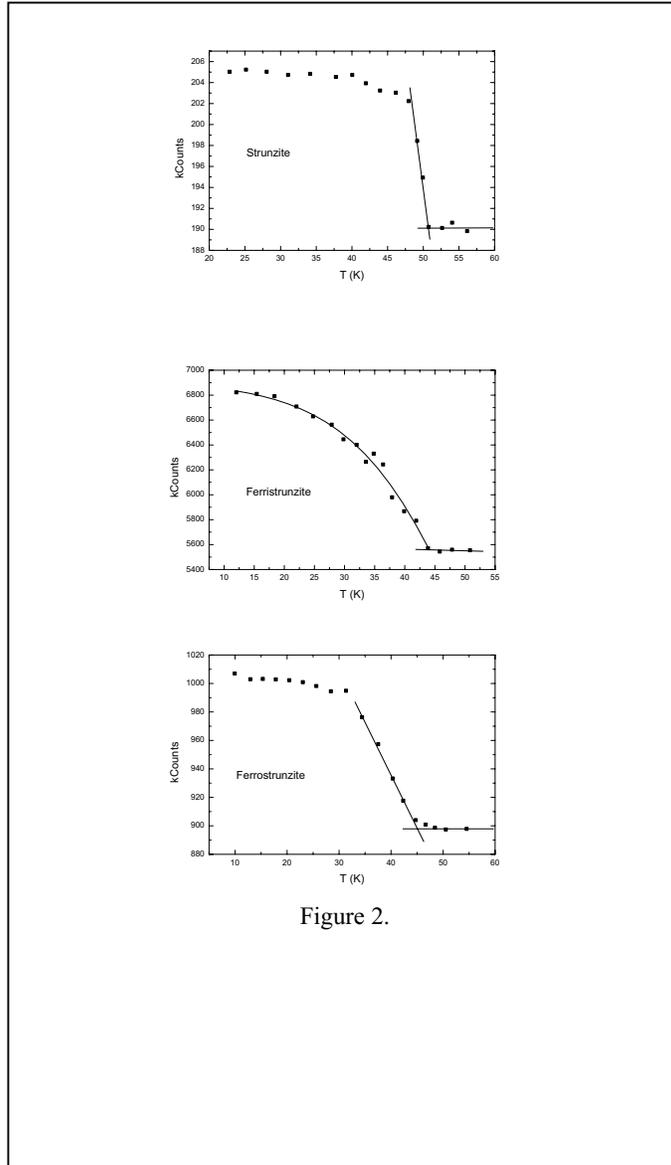


Figure 2.

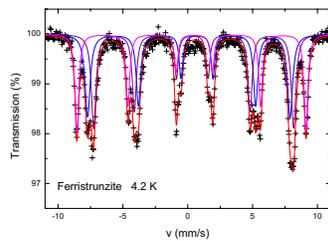


Figure 3a.

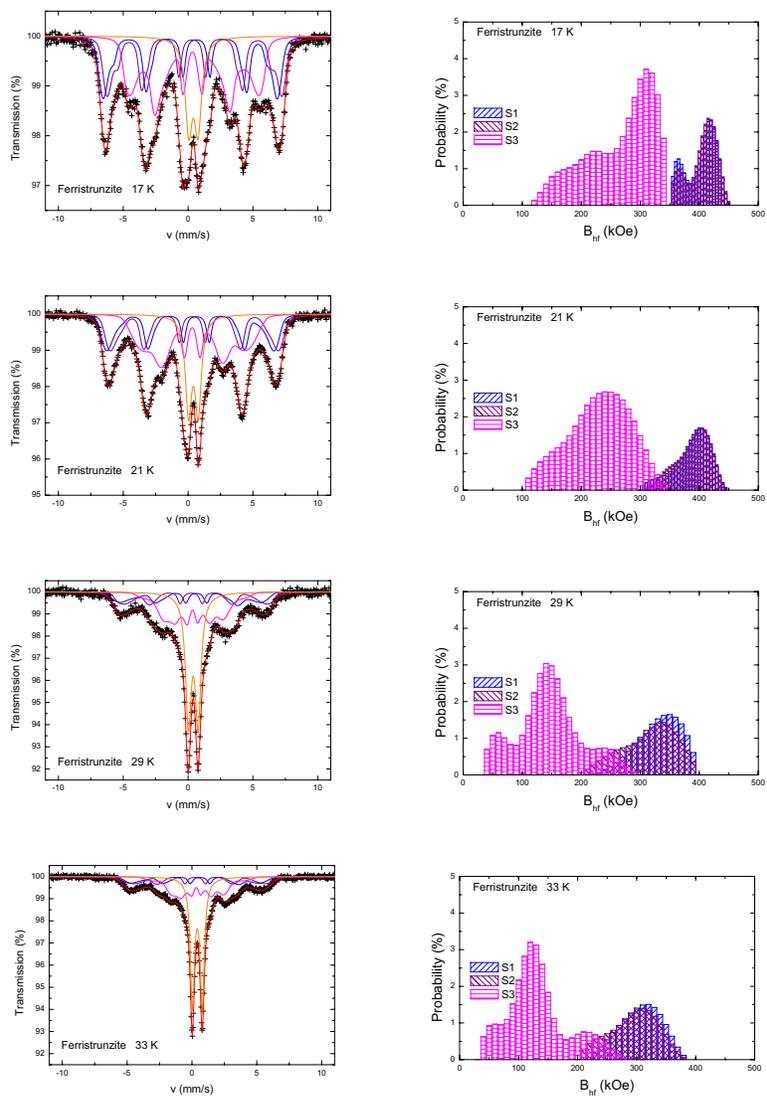


Figure 3b.

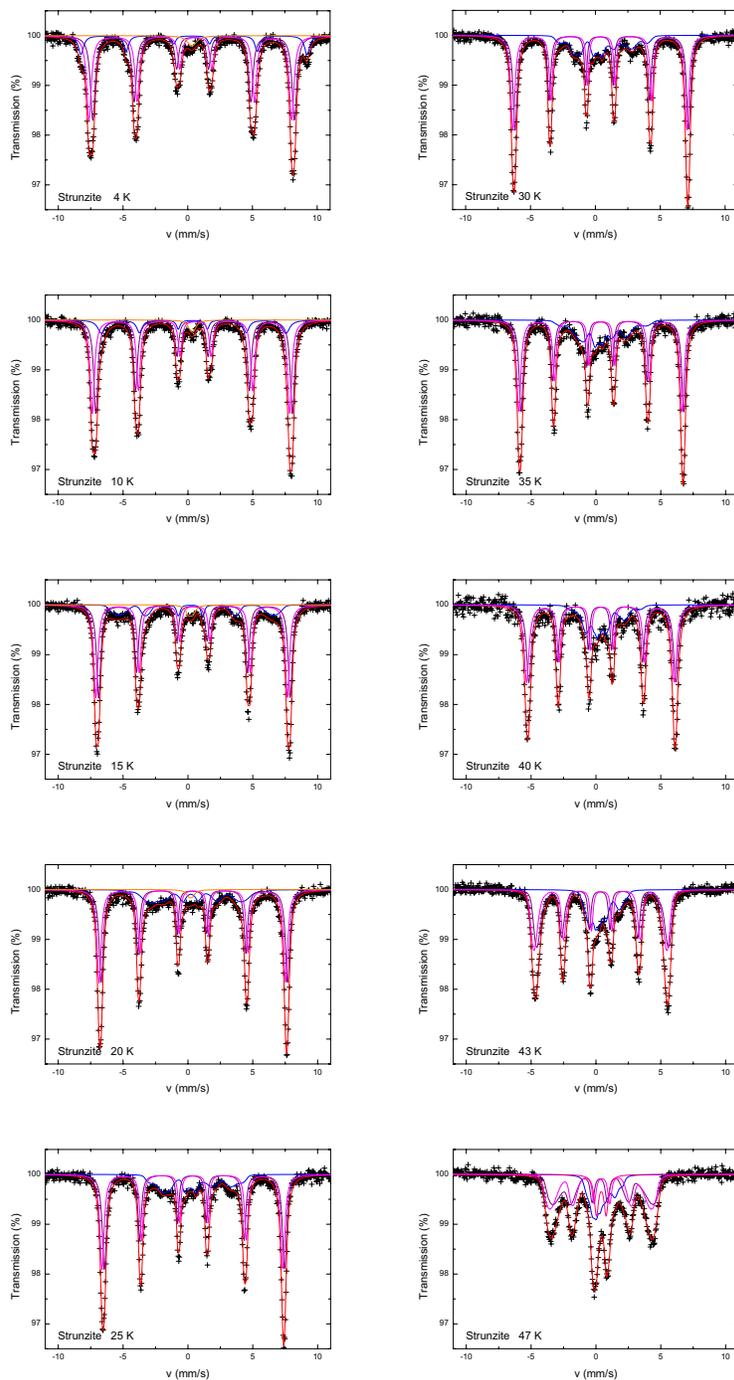


Figure 4.

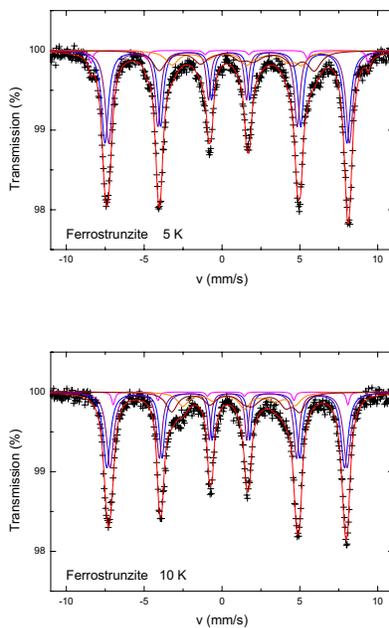


Figure 5a.

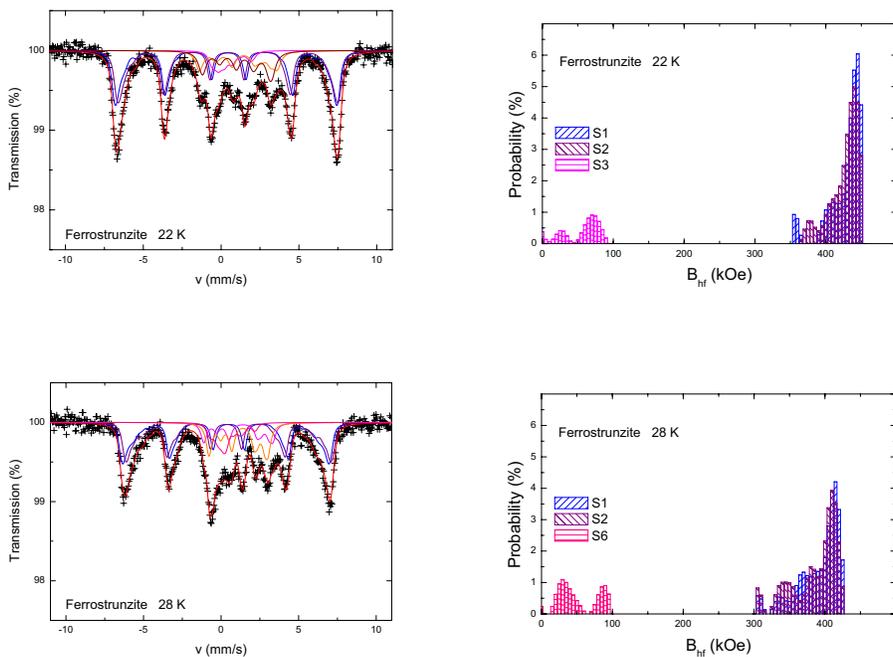


Figure 5b.

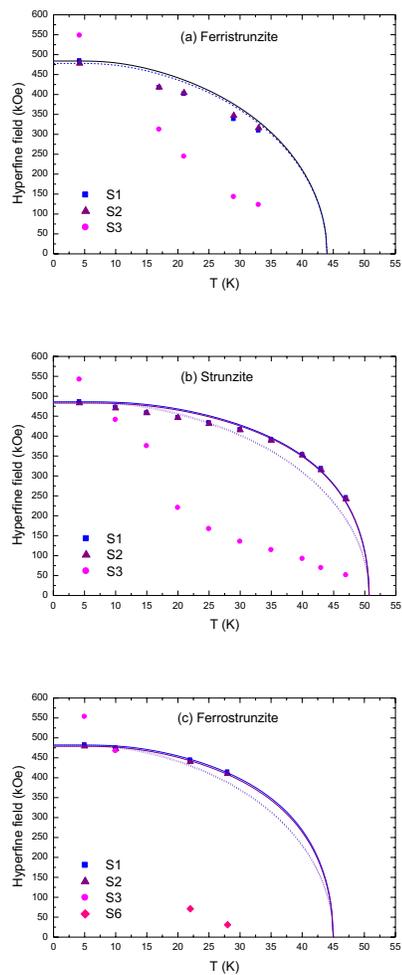


Figure 6.